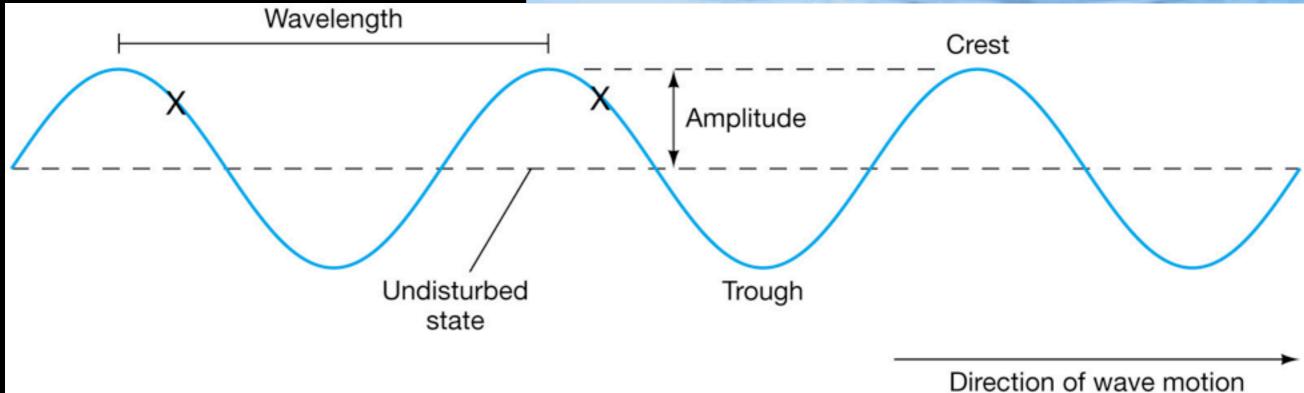
# ASTRONOMY 103: THE EVOLVING UNIVERSE

# Lecture 3 LIGHT: PARTICLES, WAVES, BLACKBODIES, AND LINES Substitute Lecturer: Paul Sell

## <u>A Short</u> <u>Description</u> <u>of Waves</u> (rope demo)





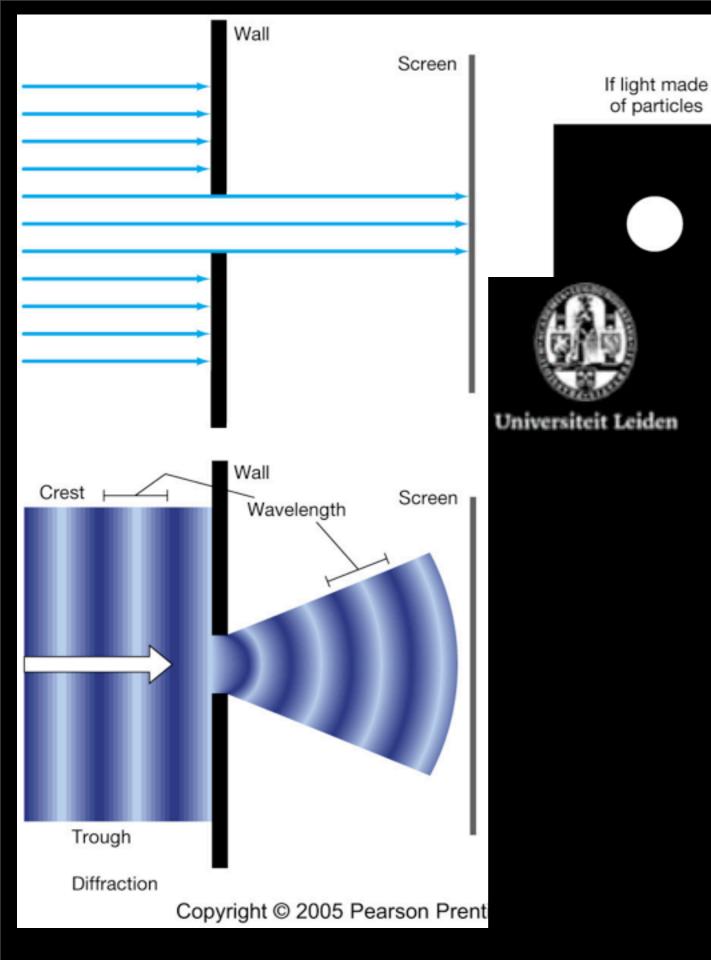
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Light is made up of photons, which are...

- a) ...waves.
- b) ...particles.
- c) ...both waves and particles.
- d) ...solid, flat beams or rays.
- e) ...all part of one big wave that is everywhere all at once.

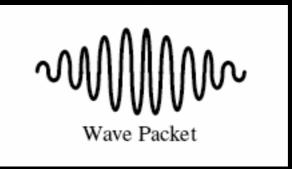
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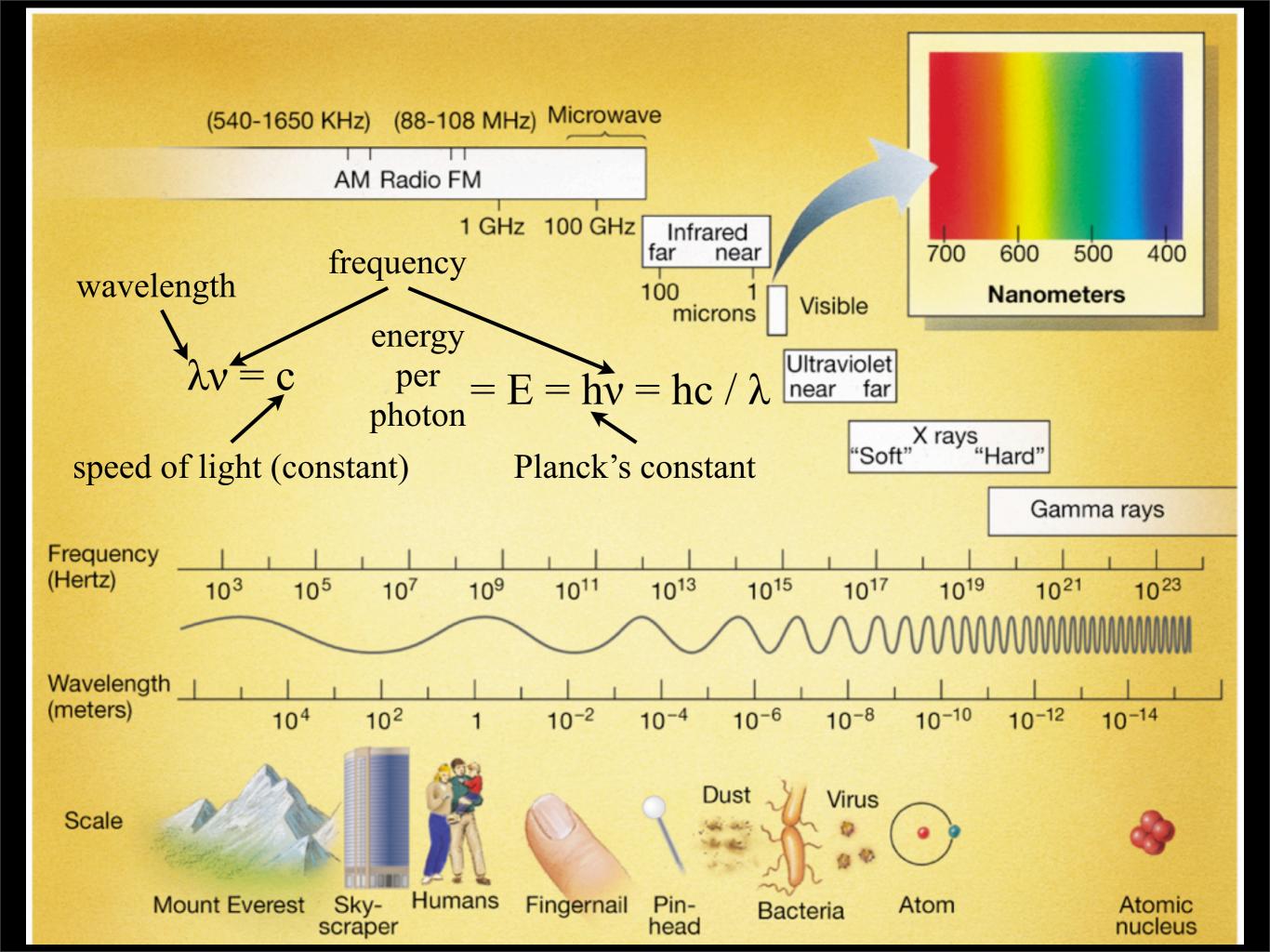


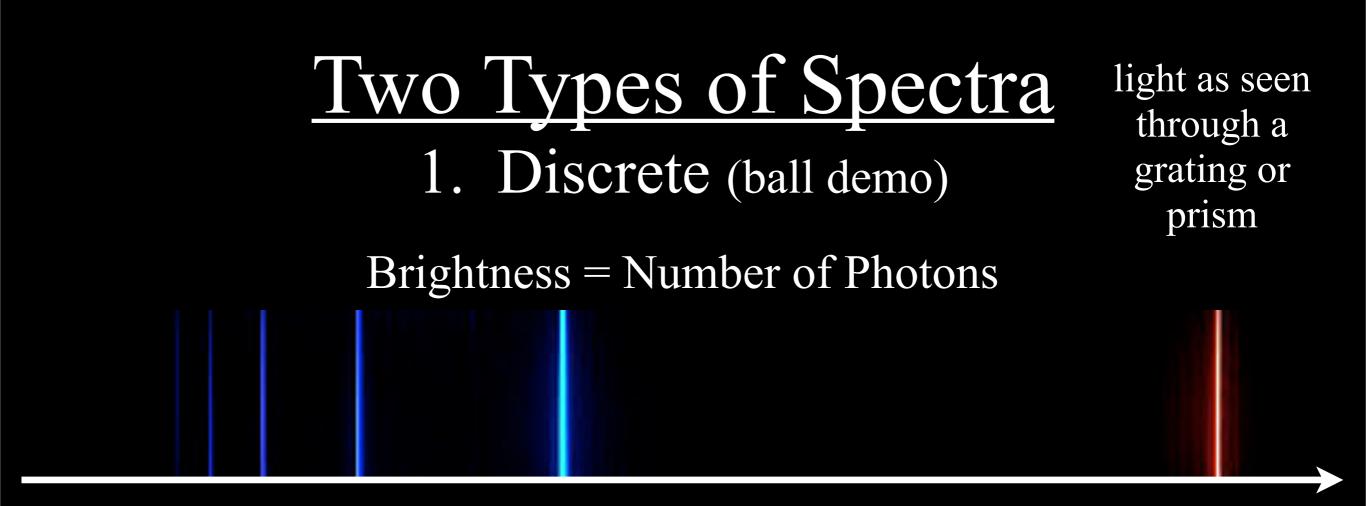
← Light comes in waves <u>Wave-Particle Duality</u> Light comes as particles

0,025 msec



A Good Compromise: Light comes in discrete electromagnetic wave packets, called photons





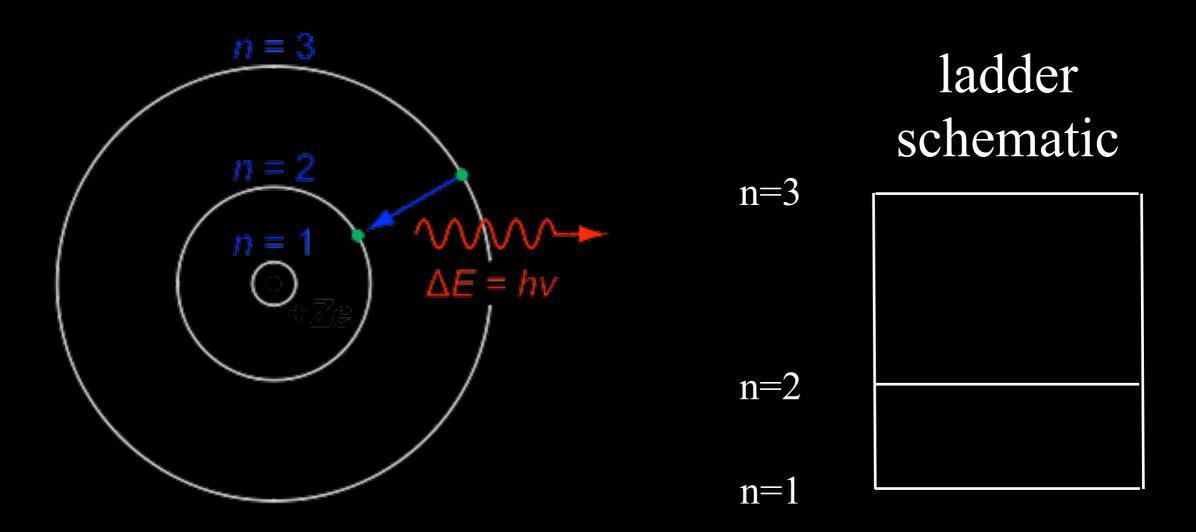
 $\lambda$  (wavelength)

Why are some lines fainter than others?

- the probability that the atom will emit a photon changes from line to line (requires quantum mechanics)
- something (e.g., dust, gas) gets in the way to block (called "extinction") or redirect (called "scattering") the light; these two processes together "attenuate" the light

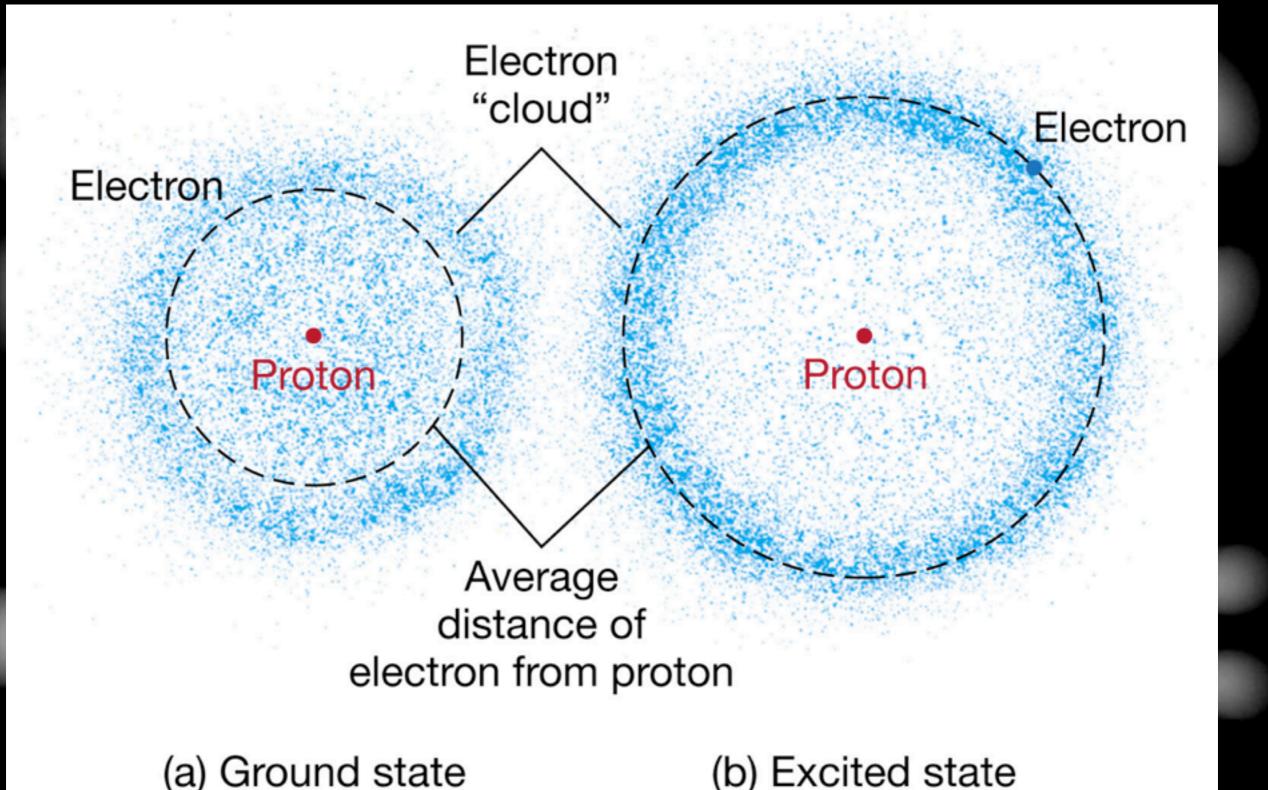
#### Two Types of Spectra 1. Discrete (ball demo)

Why do atoms only emit at discrete wavelengths?



#### The Uncertainty Principle

#### A basic result of quantum mechanics



 $\sigma_x \sigma_p \geq$ 

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 $H\alpha$  and  $H\beta$  are two lines emitted by the Hydrogen atom.  $H\alpha$  emits at a wavelength of 656.3 nm and  $H\beta$  emits at a wavelength of 486.1 nm. Which of the following statements are true?

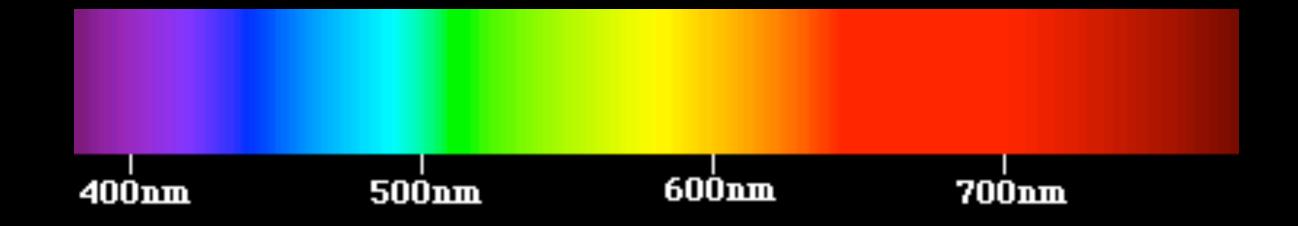
- a)  $H\beta$  emits at a lower frequency and lower energy than  $H\alpha$ .
- b)  $H\beta$  emits at a lower frequency and higher energy than  $H\alpha$ .
- c)  $H\beta$  emits at a higher frequency and lower energy than  $H\alpha$ .
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- e) There is not enough information to answer the question.

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- e) There is not enough information to answer the question.

$$\lambda v = c$$
  $E = hv = hc / \lambda$ 

#### Two Types of Spectra 2. Continuous

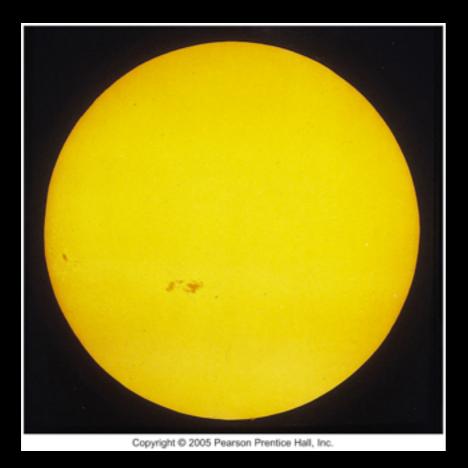


One very important type of continuous spectrum that you will see repeatedly throughout the semester: "Blackbody"

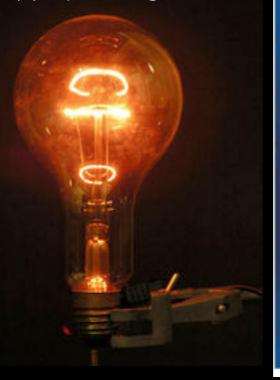
What is a blackbody?

- An ideal thermal emitter (object glows).
- A perfect absorber: no light is reflected so it appears black.
- Light comes from the heat of an opaque object.

#### Blackbody Examples:



With Ben





#### Not Blackbodies:













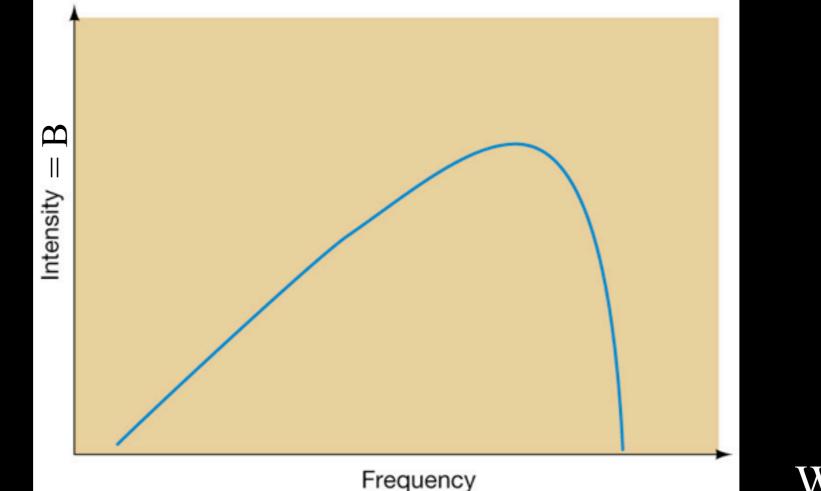




#### A Blackbody Has a Characteristic Shape

Blackbody Function = Planck Function

$$B_{\nu}(T) = \frac{2h\nu^3}{c^2} \frac{1}{e^{\frac{h\nu}{k_{\rm B}T}} - 1}; \text{ or } B_{\lambda}(T) = \frac{2hc^2}{\lambda^5} \frac{1}{e^{\frac{hc}{\lambda k_{\rm B}T}} - 1}.$$



Rayleigh-Jeans Tail:

$$B_{\nu}(T) = \frac{2\nu^2 kT}{c^2}.$$
$$B_{\lambda}(T) = \frac{2ckT}{\lambda^4},$$

Wien Tail:

$$I(\nu,T) = \frac{2h\nu^3}{c^2}e^{-\frac{h\nu}{kT}} \quad I(\lambda,T) = \frac{2hc^2}{\lambda^5}e^{-\frac{hc}{\lambda kT}}$$

#### Two Blackbody Trends

1. Wein's (Veen's) Law  $\lambda_p \propto 1 / T$ or  $\lambda_{\rm p} = 2900 \ / \ {\rm T}$  $(\lambda_p \text{ is the "peak wavelength" in } \lambda_{max} = 250 \text{ nm}$   $\lambda_{max} = 500 \text{ nm}$ micrometers or microns and T is the temperature in Kelvin)



#### 2. Stephan-Boltzman Law $F_{tot} \propto T^4$ or $F_{tot} = \sigma T^4$

(F<sub>tot</sub> is the "total flux" or the amount of photons emitted per unit area from the surface of the object)

